



Monday 24 June 2013 – Morning

GCSE TWENTY FIRST CENTURY SCIENCE PHYSICS A

A183/02 Module P7 (Higher Tier)

Candidates answer on the Question Paper. A calculator may be used for this paper.

OCR supplied materials:

None

Other materials required:

- Pencil
- Ruler (cm/mm)

Duration: 1 hour



Candidate forename				Candidate surname			
		Τ					1
Centre number				Candidate nu	umber		

INSTRUCTIONS TO CANDIDATES

- Write your name, centre number and candidate number in the boxes above. Please write clearly and in capital letters.
- Use black ink. HB pencil may be used for graphs and diagrams only.
- Answer all the questions.
- Read each question carefully. Make sure you know what you have to do before starting your answer.
- Write your answer to each question in the space provided. Additional paper may be used if necessary but you must clearly show your candidate number, centre number and question number(s).
- Do **not** write in the bar codes.

INFORMATION FOR CANDIDATES

- Your quality of written communication is assessed in questions marked with a pencil ().
- A list of useful relationships is printed on pages two and three.
- The number of marks is given in brackets [] at the end of each question or part question.
- The total number of marks for this paper is 60.
- This document consists of 16 pages. Any blank pages are indicated.



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TWENTY FIRST CENTURY SCIENCE EQUATIONS

Useful relationships

The Earth in the Universe

Sustainable energy

energy transferred = power
$$\times$$
 time
power = voltage \times current
efficiency = $\frac{\text{energy usefully transferred}}{\text{total energy supplied}} \times 100\%$

Explaining motion

$$speed = \frac{distance\ travelled}{time\ taken}$$

$$acceleration = \frac{change\ in\ velocity}{time\ taken}$$

$$momentum = mass\ \times\ velocity$$

$$change\ of\ momentum\ =\ resultant\ force\ \times\ time\ for\ which\ it\ acts$$

$$work\ done\ by\ a\ force\ =\ force\ \times\ distance\ moved\ in\ the\ direction\ of\ the\ force$$

$$amount\ of\ energy\ transferred\ =\ work\ done$$

$$change\ in\ gravitational\ potential\ energy\ =\ weight\ \times\ vertical\ height\ difference$$

$$kinetic\ energy\ =\ \frac{1}{2}\ \times\ mass\ \times\ [velocity]^2$$

Electric circuits

$$\begin{aligned} & power = voltage \times current \\ & resistance = \frac{voltage}{current} \\ & \frac{voltage \ across \ primary \ coil}{voltage \ across \ secondary \ coil} = \frac{number \ of \ turns \ in \ primary \ coil}{number \ of \ turns \ in \ secondary \ coil} \end{aligned}$$

Radioactive materials

energy = mass
$$\times$$
 [speed of light in a vacuum]²

Observing the Universe

lens power =
$$\frac{1}{\text{focal length}}$$

magnification = $\frac{\text{focal length of objective lens}}{\text{focal length of eyepiece lens}}$

speed of recession = Hubble constant × distance

pressure × volume = constant

 $\frac{\text{pressure}}{\text{temperature}}$ = constant

 $\frac{\text{volume}}{\text{temperature}}$ = constant

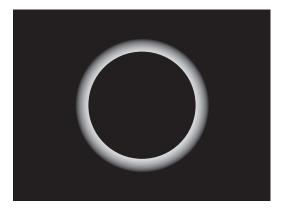
energy = mass \times [speed of light in a vacuum]²

Answer all the questions

			Al	nswer an the questi	UIIS.		
(a)	A lens w	orks by be	nding the lig	ht rays that pass th	rough it.		
	Complet	te the sente	ences that d	escribe the process	of refraction.		
	Use wor	ds from thi	s list.				
	am	plitude	frequen	cy power	speed	wavelength	
	As the li	ght enters	the glass of	the lens the		changes.	
	This res	ults in a ch	ange in				
	Howeve	r, the		cannot cha	ange.		
(b)	Here are	some dat	a on five len	282			
(2)							
	All the le	enses are n	nade from th	ne same type of glas	SS.		
		Le	ens	Diameter in cm	Focal length	in mm	
			A	10	500		
			В	20	1000		
			С	6	1000		
			D	10	20		
			E	15	35		
	obje eye	ective lens piece lens		e the best to use to			e in (i) .
(c)	Two lens	ses have ex	cactly the sa	magnification me size and shape.	=		
	Suggest	how they	can have dif	ferent focal lengths.			

One problem with using a lens is that light of different wavelengths will produce a spectrum.
Explain how the edge of a lens can produce a spectrum.
Use a diagram to help you explain.
[2]
Explain why the aperture of a radio telescope must be much larger than the aperture of an optical telescope.
[3]
[Total: 13]

2 Solar eclipses are rare.



Explain solar eclipses and why they only occur infrequently.

Include diagrams in your answer.

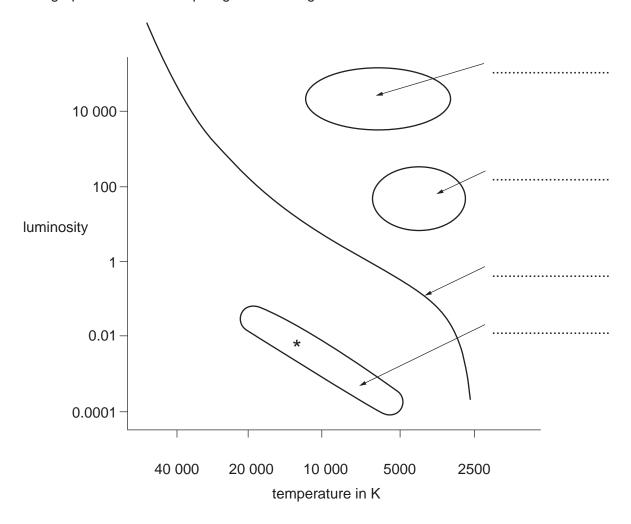


The quality of written communication will be assessed in your answer.

•••••	 	
•••••	 	
		[6]

[Total: 6]

3 The graph shows a Hertzsprung-Russell diagram.



The luminosity of the Sun is 1.

- (a) Complete the labels for the regions on the Hertzsprung-Russell diagram. [4]
- (b) Put a cross on the Hertzsprung-Russell diagram to show the position of the Sun. [1]
- (c) At the end of their lives most stars cool down and emit less and less energy.

Draw an arrow on the Hertzsprung-Russell diagram to show the direction the star (*) would move as it cools down and emits less energy. [2]

[Total: 7]

[Total: 6]

4 Scientists often disagree about how to interpret data. Additional data can often help them agree on the conclusion.

The Curtis-Shapley debate was an example of a disagreement. Additional evidence from Edwin Hubble resolved the problem.

Use this example to show how scientists can disagree about how to interpret data, until more evidence resolves the problem.

The quality of written communication will be assessed in your answer.
 [6]

[Total: 5]

5 Look at the data about some galaxies.

Galaxy location	Distance in millions of light years	Speed away from Earth in light years per year	Time taken to travel the distance at a constant speed in millions of years
Ursa Major	990	0.051	19400
Corona Borealis	1440	0.072	20 000
Bootes	2740	0.131	
Hydra	3960	0.198	

(a)	Complete the table. Assume that the speeds of each of the galaxies is constant.	[1]
(b)	14 000 million years ago?	-
		[4]

Jo i	s researching a red giant star that is app	proximately the same mass as the Sun.	
(a)	Jo knows the surface temperature of th	e red giant is approximately 3400 K.	
	What is this temperature in °C?		
		20	
	τε	emperature =°C	, [2]
(b)	In a red giant the main source of energ	y is the fusion of helium nuclei in the core .	
	Explain how this shows that the core to core temperature of the Sun.	temperature of the red giant must be higher than	the
			[2]
(c)	Which of the following elements are no	t produced by fusion in this low mass red giant?	
	Put ticks (✓) in the boxes next to the co	orrect answers.	
	carbon		
	hydrogen		
	iron		
	nitrogen		
	oxygen		[2]
			[4]

6

(d)	What is likely to be the next stage in	n the life of this low mass red giant?	
	Put a tick (\checkmark) in the box next to the	correct answer.	
	blue giant		
	main sequence		
	protostar		
	supernova		
	white dwarf		[1]
			[Total: 7]

[Total: 6]

7 Read this extract from an article.

A telescope in Arizona has taken the clearest pictures ever taken from an Earth-based telescope. The images are three times sharper than those from the Hubble Space Telescope. The telescope uses adaptive optics.

Adaptive optics compensate for some of the effects of the atmosphere by gently bending an ultrathin secondary mirror that corrects the incoming light. This mirror is so thin that it can be bent into different shapes, as controlled by sensors that detect atmospheric distortions.

Suggest how the new adaptive optics affect the balance of advantages and disadvantages of ground-based and space-based telescopes.

The quality of written communication will be assessed in your answer.
[6]

			10
8	At t	he be	eginning of the 20th century, scientists could not explain how the Sun produced its energy.
			r, the development of Einstein's equation $E=mc^2$ showed that energy could be released ass is lost during nuclear fusion, where $c=3 \times 10^8 \text{m/s}$.
	(a)	The	luminosity of the Sun is about 3.90 x 10 ²⁶ J/s.
		(i)	Use Einstein's equation to calculate the amount of mass lost each second by the Sun.
			mass loss per second = kg/s [3]
		(ii)	The Sun will be on the main sequence for about 10 billion (10 ¹⁰) years.
			Assume that the only loss of mass from the Sun is due to the fusion of hydrogen.
			How much mass will the Sun lose while it is on the main sequence?
			mass loss =
	(b)		st of the energy produced in the Sun is from the fusion of hydrogen into helium by the con-proton chain.
		The	proton-proton chain involves three stages.
			${}_{1}^{1}H + {}_{1}^{1}H \rightarrow {}_{1}^{2}H + e^{+}$
			${}_{1}^{2}H + {}_{1}^{1}H \rightarrow {}_{2}^{3}He$
			${}_{2}^{3}\text{He} + {}_{2}^{3}\text{He} \rightarrow {}_{2}^{4}\text{He} + {}_{1}^{1}\text{H} + {}_{1}^{1}\text{H}$
		(i)	How many positrons (e ⁺) are emitted to produce a stable ⁴ He nucleus?
			[1]
		(ii)	Why is a positron emitted in the first stage and not an electron?
			[1]
			[Total: 7]

9	(a)	This question is about naked eye astronomy.	
		Which time is the shortest?	
		Put a tick (✓) in the box next to the correct answer.	
		The time taken for the Moon to return to the same position in the sky.	
		24 hours.	
		The time taken for a star to return to the same position in the sky.	
		The time for the Sun to return to its highest position in the sky.	[1]
	(b)	How do astronomers describe the position of the Sun, Moon and stars in the sky?	
		Γ	Total: 3]

END OF QUESTION PAPER

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