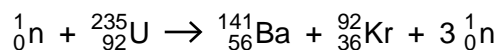


- 1 (a) The following nuclear reaction occurs when a slow-moving neutron is absorbed by an isotope of uranium-235.



- (i) Explain how this reaction is able to produce energy.

.....

 [2]

- (ii) State in what form the energy is released in such a reaction.

..... [1]

- (b) The binding energy per nucleon of each isotope in (a) is given in Fig. 8.1.

isotope	binding energy per nucleon/MeV
${}_{92}^{235}\text{U}$	7.6
${}_{56}^{141}\text{Ba}$	8.3
${}_{36}^{92}\text{Kr}$	8.7

Fig. 8.1

- (i) Explain why the neutron ${}_0^1\text{n}$ does not appear in the table above.

.....
 [1]

- (ii) Calculate the energy released in the reaction shown in (a).

energy = MeV [2]

[Total: 6]

2 A proton travelling at a high velocity is fired at a stationary proton. It stops momentarily at a distance of 2.0×10^{-15} m from the stationary proton.

(a) Calculate the electrostatic force acting on each proton when separated by 2.0×10^{-15} m.

force = N [2]

(b) The two protons fuse together. Explain how the protons are able to remain together.

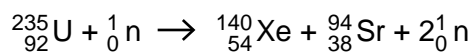
.....
..... [1]

(c) Explain why the proton must have a very large velocity for the fusion to occur and the protons to remain together.

.....
.....
.....
..... [2]

[Total: 5]

- 3 (a) In the core of a nuclear reactor, one of the many fission reactions of the uranium-235 nucleus is shown below.



- (i) State **one** quantity that is conserved in this fission reaction.

..... [1]

- (ii) Fig. 4.1 illustrates this fission reaction.

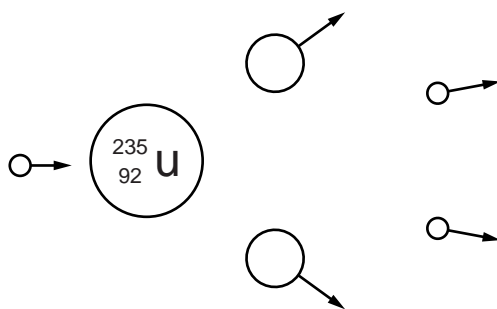
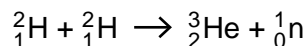


Fig. 4.1

Label all the particles in Fig. 4.1 and extend the diagram to show how a chain reaction might develop. [2]

- (b) Fusion of hydrogen nuclei is the source of energy in most stars. A typical reaction is shown below.



The ${}^2_1\text{H}$ nuclei repel each other. Fusion requires the ${}^2_1\text{H}$ nuclei to get very close and this usually occurs at very high temperatures, typically 10^9K .

(i) Use the data below to calculate the energy released in the fusion reaction above.

mass of ${}^2_1\text{H}$ nucleus = 3.343×10^{-27} kg

mass of ${}^3_2\text{He}$ nucleus = 5.006×10^{-27} kg

mass of ${}^1_0\text{n}$ = 1.675×10^{-27} kg

energy = J [3]

(ii) State in what form the energy in (b)(i) is released.

..... [1]

(iii) The ${}^2_1\text{H}$ nuclei in stars can be modelled as an ideal gas. Calculate the mean kinetic energy of the ${}^2_1\text{H}$ nuclei at 10^9 K.

energy = J [2]

(iv) Suggest why some fusion can occur at a temperature as low as 10^7 K.

.....
.....
..... [1]

[Total: 10]

- 4 The isotopes of carbon-14 ($^{14}_6\text{C}$) and carbon-15 ($^{15}_6\text{C}$) are beta-minus emitters. The table in Fig. 5.1 shows the maximum kinetic energy of each electron emitted and the half-life of the isotope.

isotope	maximum kinetic energy / MeV	half-life
$^{14}_6\text{C}$	0.16	5560 years
$^{15}_6\text{C}$	9.8	2.3 s

Fig. 5.1

- (a) State one property common to all isotopes of an element.

.....
 [1]

- (b) The neutrons and protons inside each isotope experience fundamental forces. Name the two fundamental forces experienced by both neutrons and protons.

1.
 2. [2]

- (c) An isotope of carbon-15 decays into an isotope of nitrogen (N).

- (i) Complete the nuclear reaction below.



- (ii) Use the quark model to state the changes taking place within the nucleus of the carbon-15 atom.

.....
 [1]

- (d) (i) Estimate the maximum speed of an electron from the nucleus of carbon-14.

speed = ms^{-1} [2]

(ii) Suggest why the actual speed of the electron is much less than your answer in (i).

.....
..... [1]

(e) (i) Calculate the decay constant λ in s^{-1} of carbon-14.

$\lambda = \dots\dots\dots \text{s}^{-1}$ [2]

(ii) The molar mass of carbon-14 is 14 g mol^{-1} . Show that 1.0 mg of carbon-14 has 4.3×10^{19} nuclei.

[1]

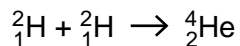
(iii) Calculate the activity of the 1.0 mg mass of carbon-14.

activity = $\dots\dots\dots \text{Bq}$ [2]

5 (a) Explain the term *binding energy* of a nucleus.

.....
.....
..... [2]

(b) Nuclear fusion takes place in the core of the Sun. One of the simplest fusion reactions is shown below.



(i) The binding energy per nucleon of ${}^2_1\text{H}$ is $1.8 \times 10^{-13} \text{ J}$ and the binding energy per nucleon of ${}^4_2\text{He}$ is $1.1 \times 10^{-12} \text{ J}$. Show that the energy released in the reaction is $3.7 \times 10^{-12} \text{ J}$.

(ii) The Sun radiates its energy uniformly through space. The mean intensity of the Sun's radiation reaching the Earth's atmosphere is about 1400 W m^{-2} . The mean radius of the Earth's orbit round the Sun is $1.5 \times 10^{11} \text{ m}$.

1 Show that the mean power radiated from the surface of the Sun is $4.0 \times 10^{26} \text{ W}$.

[2]

2 Assume all the radiated energy from the Sun comes from the fusion reaction shown in (b). Estimate the number of helium-4 nuclei produced every second by the Sun.

number = s^{-1} [2]

[Total: 8]