

Question Number	Answer	Mark
<b>1(a)(i)</b>	Reverse direction for temperature [at least 2 values seen] (1)	<b>2</b>
	Logarithmic/power temperature variation [at least 3 realistic values seen increasing by the same factor] (1)	
<b>1(a)(ii)</b>	<p>QWC – Work must be clear and organised in a logical manner using technical wording where appropriate</p> <p><b>Area 1: Max 2</b></p> <p>The Sun is fusing/burning hydrogen (into helium in its core) (1)</p> <p>When (hydrogen) fusion/burning ceases the core of the Sun cools [accept radiation pressure drops when fusion/burning ceases in the core] (1)</p> <p>The core collapses/contracts (under gravitational forces) (1)</p> <p><b>Area 2: Max 2</b> (1)</p> <p>The Sun expands and becomes a red giant (1)</p> <p>The core becomes hot enough for helium fusion/burning to begin (in the core) (1)</p> <p>Helium begins to run out and the core collapses again (under gravitational forces) (1)</p> <p><b>Area 3: Max 2</b> (1)</p> <p>Idea that outer layers of Sun are ejected into space (1)</p> <p>The temperature doesn't rise enough for further fusion to begin (1)</p> <p>The core/Sun becomes a (white) dwarf star</p>	<b>6</b>
<b>1(b)(i)</b>	Idea of a very high temperature [accept value of about $10^7$ K] (1)	<b>3</b>
	<p>To overcome repulsive/electrostatic forces between protons/nuclei  <b>Or</b> so that protons/nuclei get close enough together for the strong (nuclear) force to act  <b>Or</b> so that protons/nuclei get close enough to fuse (1)</p> <p>Idea of a very high density [accept pressure] to give a sufficient collision rate (1)</p>	
<b>1(b)(ii)</b>	<p>Attempt at calculation of mass deficit (1)</p> <p>Use of <math>\Delta E = c^2 \Delta m</math> (1)</p> <p>Attempt at conversion from J to (M)eV (1)</p> <p><math>\Delta E = 12.9</math> (MeV) (1)</p> <p>[If correct mass defect in kg is converted into u and then <math>1u = 931</math> Mev used, then full marks may be awarded]</p> <p><u>Example of calculation</u></p> <p>Physics And Math (50808802) – <math>6.646483 - (1.673534 \times 2) \times 10^{-27}</math> kg</p>	<b>4</b>

	$\Delta m = 2.2925 \times 10^{-29} \text{ kg}$ $\Delta E = (3.00 \times 10^8 \text{ ms}^{-1})^2 \times 2.2925 \times 10^{-29} \text{ kg} = 2.063 \times 10^{-12} \text{ J}$ $\Delta E = \frac{2.063 \times 10^{-12} \text{ J}}{1.60 \times 10^{-13} \text{ JMeV}^{-1}} = 12.9 \text{ MeV}$	
	<b>Total for question</b>	<b>15</b>

Question Number	Answer	Mark															
2(a)	<b>Similarity:</b> Same number of protons <b>Or</b> same magnitude of charge <b>Or</b> both have 1 proton (1)	2															
	<b>Difference:</b> Different number of neutrons / nucleons <b>Or</b> different mass <b>Or</b> D has 1 neutrons and T has 2 neutrons (1)																
2(b)	Use of $P = \frac{\Delta E}{\Delta t}$ (do not penalise a power of ten error) (1)	2															
	Energy = $7.5 \times 10^6$ (J) (1)  <u>Example of calculation</u> $E = 500 \times 10^{12} \text{ W} \times 15 \times 10^{-9} \text{ s} = 7.5 \times 10^6 \text{ J}$																
2(c)(i)	${}^2_1\text{D} + {}^3_1\text{T} \rightarrow {}^4_2\text{He} + {}^1_0\text{n}$	2															
	<table border="1"> <tr> <td>Top line</td> <td>2</td> <td>3</td> <td>4</td> <td>1</td> </tr> <tr> <td>Bottom line</td> <td>1</td> <td>1</td> <td>2</td> <td>0</td> </tr> </table> (1) (1)		Top line	2	3	4	1	Bottom line	1	1	2	0					
Top line	2	3	4	1													
Bottom line	1	1	2	0													
2(c)(ii)	Attempt at calculation of mass difference (1) Energy released = 17.5 (MeV) [17.5 must be clearly identified as an energy] (1)	2															
	<u>Example of calculation</u> $\Delta m = (1875.6 + 2808.9 - 3727.4 - 939.6) \text{ MeV}/c^2 = 17.5 \text{ MeV}/c^2$ $\Delta E = 17.5 \text{ MeV}$																
2(c)(iii)	Conversion of energy to consistent units (1) Number of nuclei = $3 \times 10^{18}$ (1)	2															
	<u>Example of calculation</u> In each fusion $\Delta E = 17.5 \times 10^6 \text{ eV} \times 1.6 \times 10^{-19} \text{ J eV}^{-1} = 2.8 \times 10^{-12} \text{ J}$ $\therefore N = \frac{7.5 \times 10^6 \text{ J}}{2.8 \times 10^{-12} \text{ J}} = 2.68 \times 10^{18}$																
	<table border="1"> <thead> <tr> <th>Energy MJ (b)</th> <th>Energy MeV (c)(ii)</th> <th>N <math>\times 10^{18}</math></th> </tr> </thead> <tbody> <tr> <td>7.5</td> <td>17.5</td> <td>2.7</td> </tr> <tr> <td>7.5</td> <td>20</td> <td>2.3</td> </tr> <tr> <td>8</td> <td>17.</td> <td>2.9</td> </tr> <tr> <td>8</td> <td>20</td> <td>2.5</td> </tr> </tbody> </table>	Energy MJ (b)	Energy MeV (c)(ii)	N $\times 10^{18}$	7.5	17.5	2.7	7.5	20	2.3	8	17.	2.9	8	20	2.5	
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8	20	2.5															

<p><b>2(c)(iv)</b></p>	<p>Application of momentum conservation (1)</p> <p>Deduction that <math>V_N = 4 V_\alpha</math> [<math>v_N = 3.967 v_\alpha</math>] (1)</p> <p>Use of <math>E_k = \frac{1}{2}mv^2</math> (ratio as shown <b>or</b> sum = 17.5 MeV) (1)</p> <p>Energy = 14 MeV (ecf (c)(ii), 14.1 MeV, if <math>v_N = 3.967 v_\alpha</math> 16 MeV if 20 MeV used) (1)</p> <p><b>Or</b></p> <p>Application of momentum conservation (1)</p> <p>Use of <math>E_k = p^2/2m</math> (1)</p> <p>Deduction that <math>E_N = 4 E_\alpha</math> (1)</p> <p>Energy = 14 MeV (1)</p> <p><u>Example of calculation (1<sup>st</sup> method)</u></p> $m_N V_N = m_\alpha V_\alpha$ $V_N = \frac{m_\alpha}{m_N} \times V_\alpha = 4V_\alpha$ $\frac{E_N}{E_\alpha} = \frac{\frac{1}{2}m_N V_N^2}{\frac{1}{2}m_\alpha V_\alpha^2} = \frac{1}{4} \times \left(\frac{4}{1}\right)^2 = 4$ $\therefore E_N = \frac{4}{5} \times 17.5 \text{ MeV} = 14 \text{ MeV}$ <p><u>Example of calculation (2<sup>nd</sup> method)</u></p> $p_\alpha = p_N$ $p_\alpha^2 = p_N^2$ $E_\alpha \times 2m_\alpha = E_N \times 2m_N$ $\therefore E_\alpha = E_N \times \frac{m_N}{m_\alpha} = \frac{E_N}{4}$ <p>Also, <math>E_\alpha + E_N = 17.5 \text{ MeV}</math></p> $\therefore \frac{E_N}{4} + E_N = 17.5 \text{ MeV}$ $\therefore E_N = \frac{4}{5} \times 17.5 \text{ MeV} = 14 \text{ MeV}$	<p style="text-align: center;">4</p>
<p><b>2(d)</b></p>	<p><b>Max 3</b></p> <p>A heavy <b>nucleus</b> absorbs a neutron. [accepts “collides with” / “fired into” for “absorbs”] (1)</p> <p>The <b>nucleus</b> becomes unstable <b>and</b> splits into two (roughly equal sized) fragments [accept “decays” / “breaks up” for “splits”] (1)</p> <p>Idea that a few neutrons are also emitted in the fission process (1)</p> <p>These neutrons cause further fissions <b>Or</b> these neutrons cause a chain reaction (1)</p> <p>(if atom is used instead of nucleus <b>only penalise once</b>)</p>	<p style="text-align: center;">3</p>
	<p><b>Total for question</b></p>	<p style="text-align: center;">17</p>



<p><b>*3(b)(i)</b></p>	<p>(QWC- Work must be clear and organised in a logical manner using technical wording where appropriate.)</p> <p>Max THREE from first 5 marking points</p> <ul style="list-style-type: none"> <li>• Very high temperatures (<math>&gt;10^7</math> K) needed (1)</li> <li>• To overcome electrostatic repulsion / forces (1)</li> <li>• <u>Nuclei</u> come close enough to fuse / for strong (nuclear) force to act (1)</li> <li>• Very high densities needed (1)</li> <li>• (Together with high nuclei speeds) this gives a sufficient collision rate (1)</li>   <li>• (Very high) temperatures lead to confinement problems (1)</li> <li>• Contact with container causes temperature to fall (and fusion to cease) (1)</li> </ul>	<p><b>Max 4</b></p>
<p><b>3(b)(ii)</b></p>	${}^2_1D + {}^2_1D \rightarrow {}^3_1H + {}^1_1X$ <p>X is a proton [accept hydrogen nucleus]</p>	<p>(1) <b>1</b></p>
<p><b>3(b)(iii)</b></p>	<p>Any TWO from</p> <ul style="list-style-type: none"> <li>• (Hydrogen) fuel for fusion is (virtually) unlimited whereas fission relies upon (uranium) a relatively limited resource (1)</li> <li>• Fusion results in few radioactive products, but radioactive products produced in fission present significant disposal problems (1)</li> <li>• For a given mass of fuel, the energy released by fusion is greater than the energy released by fission (1)</li> </ul>	<p><b>Max 2</b></p>
<p><b>Total for question</b></p>		<p><b>15</b></p>

Question Number	Answer	Mark
4(a)	Alpha-radiation only has a range of a few cm in air / cannot penetrate walls of container / skin (1)	(1)
4(b)(i)	Top line: ${}^{241}\text{Am} \rightarrow {}^{237}\text{Np} + {}^4\alpha$ (1)	(2)
	Bottom line: ${}_{95}\text{Am} \rightarrow {}_{93}\text{Np} + {}_2\alpha$ (1)	
4(b)(ii)	Attempt at calculation of mass defect (1)	(3)
	Use of $(\Delta)E=c^2(\Delta)m$ OR use of $1 \text{ u} = 931.5 \text{ MeV}$ (1)	
	Correct answer [5.65 MeV; accept 5.6 - 5.7 MeV] (1)	
	Example of calculation: $\Delta m = 241.056822 \text{ u} - 237.048166 \text{ u} - 4.002603 \text{ u} = 0.006053 \text{ u}$ $\Delta m = 0.006053 \text{ u} \times 1.66 \times 10^{-27} \text{ kg u}^{-1} = 1.005 \times 10^{-29} \text{ kg}$ $E = 1.005 \times 10^{-29} \text{ kg} \times (3 \times 10^8 \text{ ms}^{-1})^2 = 9.04 \times 10^{-13} \text{ J}$ $E = \frac{9.04 \times 10^{-13} \text{ J}}{1.6 \times 10^{-13} \text{ MeV J}^{-1}} = 5.65 \text{ MeV}$	
4(c)	Reference to half-life and typical lifespan (1)	
	<b>Total for question</b>	<b>(7)</b>